

A detailed visualization of the cosmic web, showing a dense network of filaments and nodes of matter in deep purple and gold colors.

# Quantum Sensors for Particle Detection

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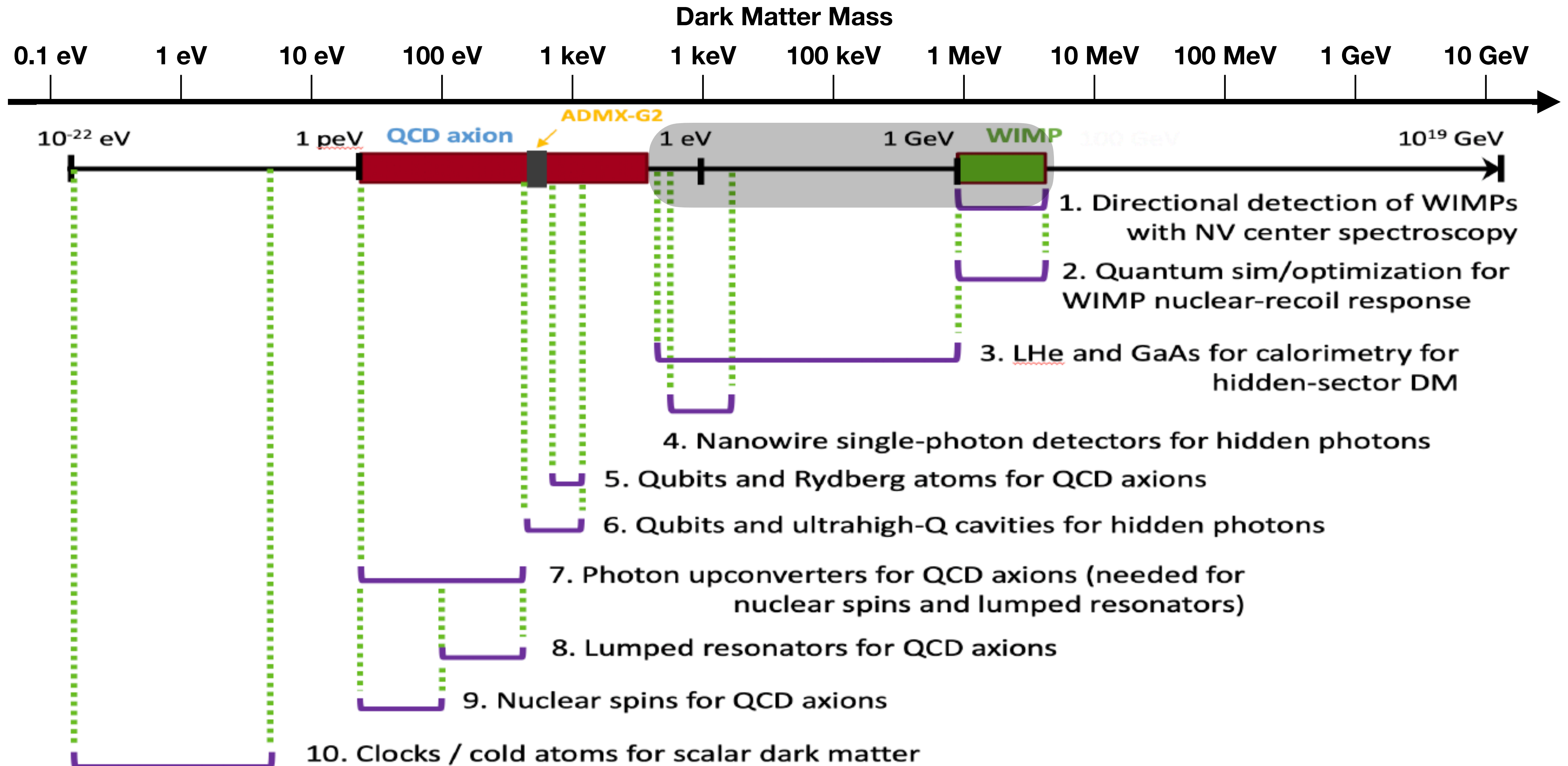


# “Quantum” and “Particle” in this talk

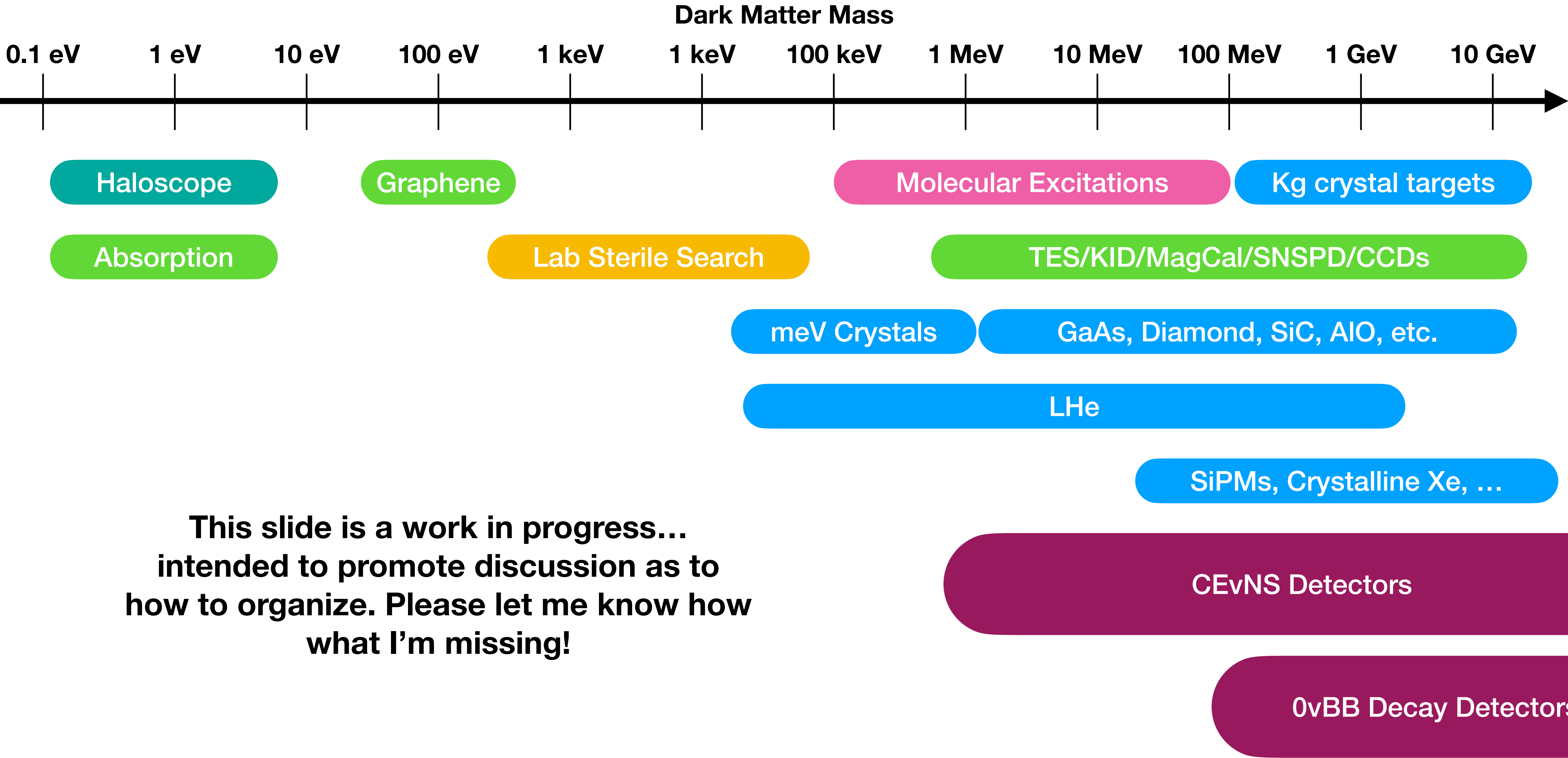
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- “Quantum” refers to techniques and technologies used in QIS work and being applied to particle detection, but also to techniques measuring single quanta, or signals from single quanta.
- “Particle” refers to localized, particle-like interactions as opposed to coherent mode sensing techniques

# Dark Matter and Neutrino Particle Detection



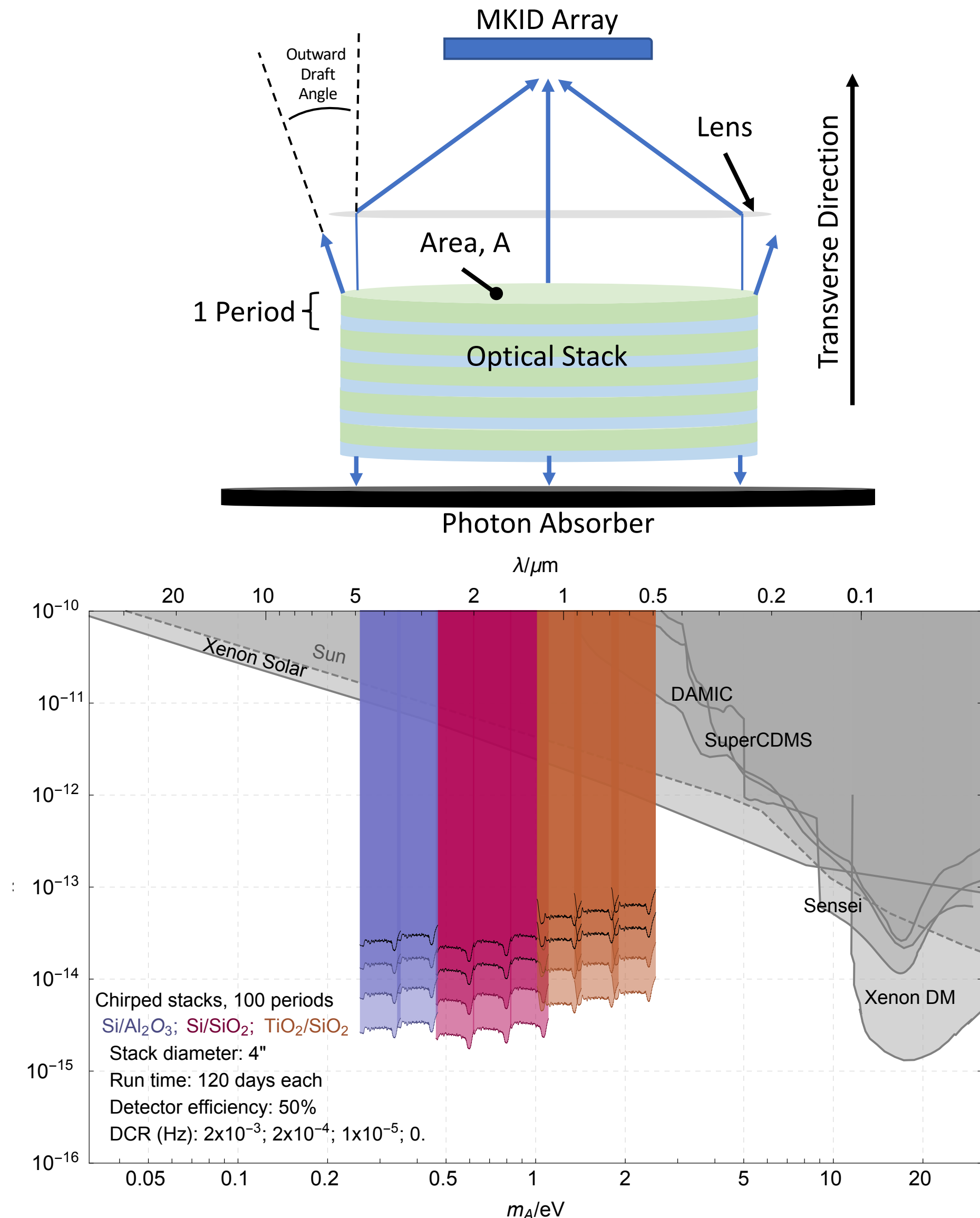
# Dark Matter and Neutrino Particle Detection





# Haloscopes

- Use stacks of different dielectric layers with alternating index of refraction to induce dark photon  $\rightarrow$  photon conversion
- photons are then focused on to a detector
- SNSPD, MKIDs, or other photon detectors could be used

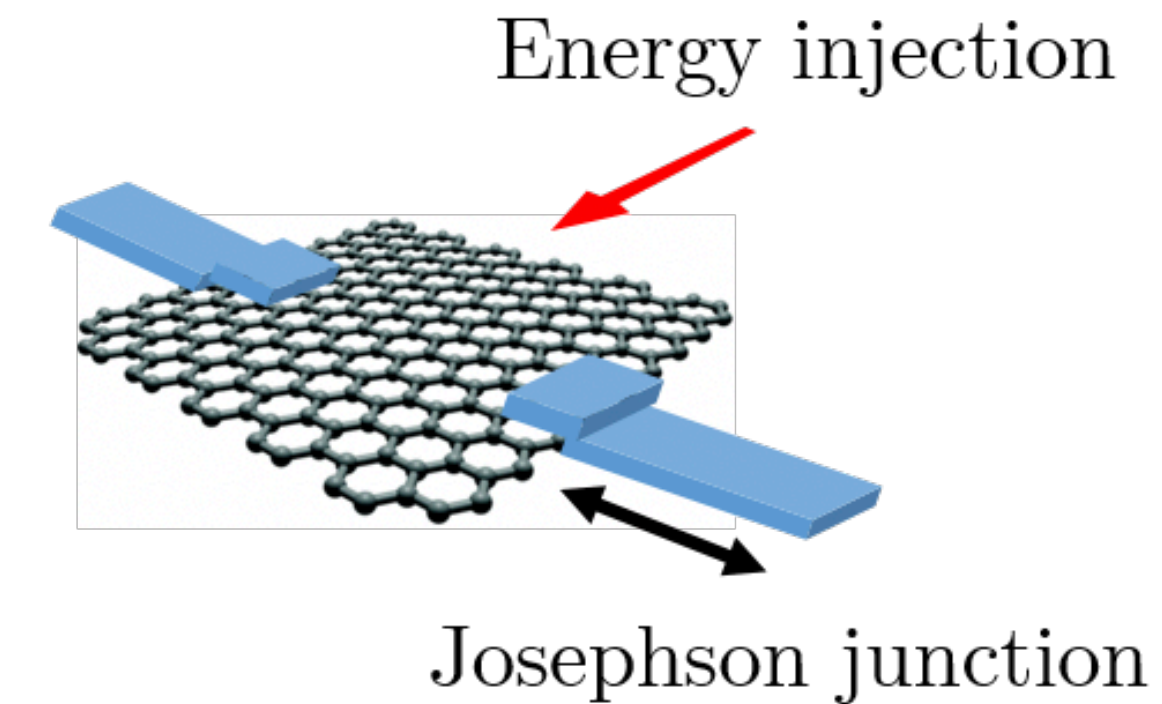




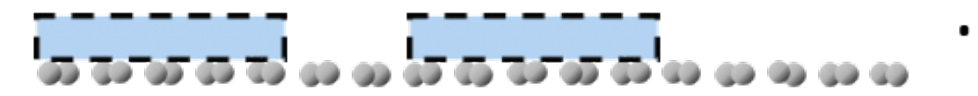
# Graphene Josephson Junctions

- Monolayer graphene acts as the normal layer in a SNS Josephson Junction.
- Injected energy is sensed by the change in the JJ from its zero-voltage to resistive state.
- Sensitivities equivalent to 0.13 meV have been demonstrated through NEP bolometric measurements.
- Sensitivity to electron recoils from 0.1 keV dark matter possible

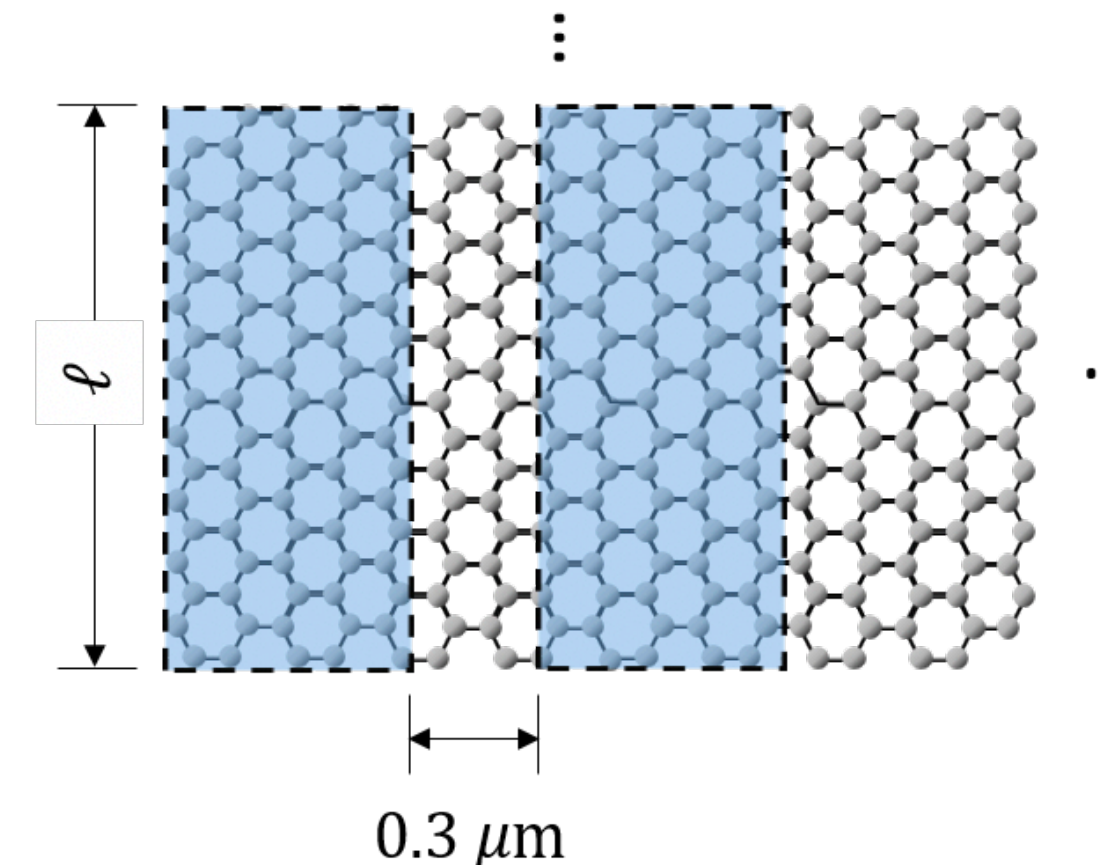
(a) Single GJJ unit



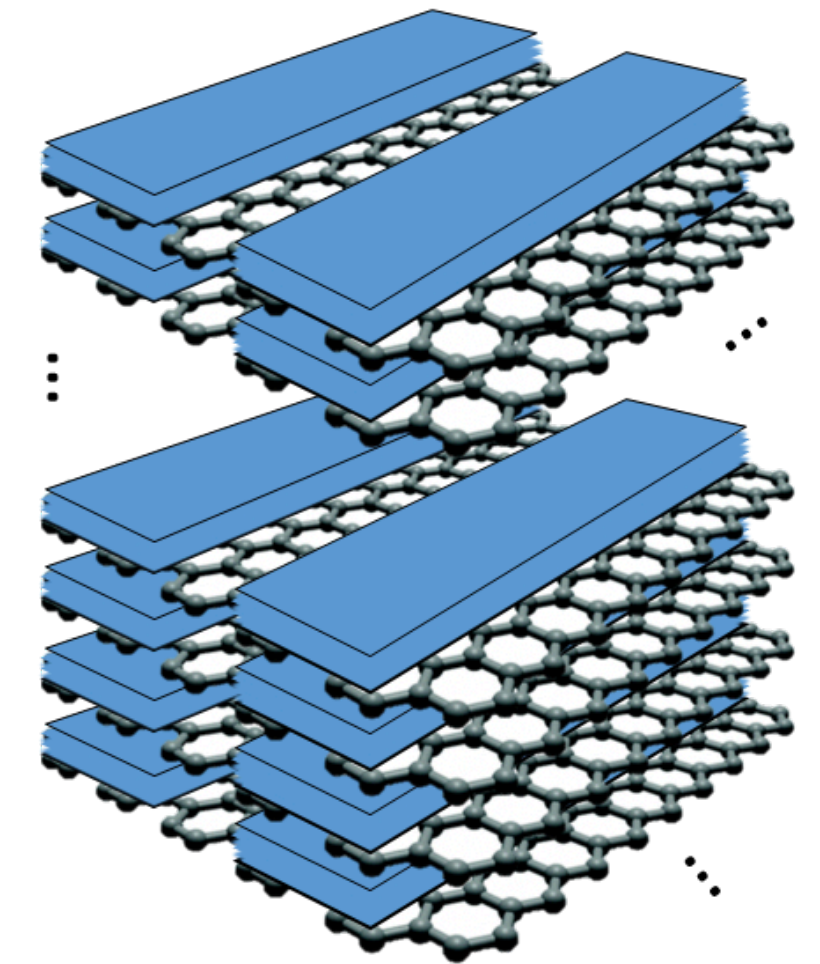
(b) Single unit: lateral view



(c) Single unit: top view



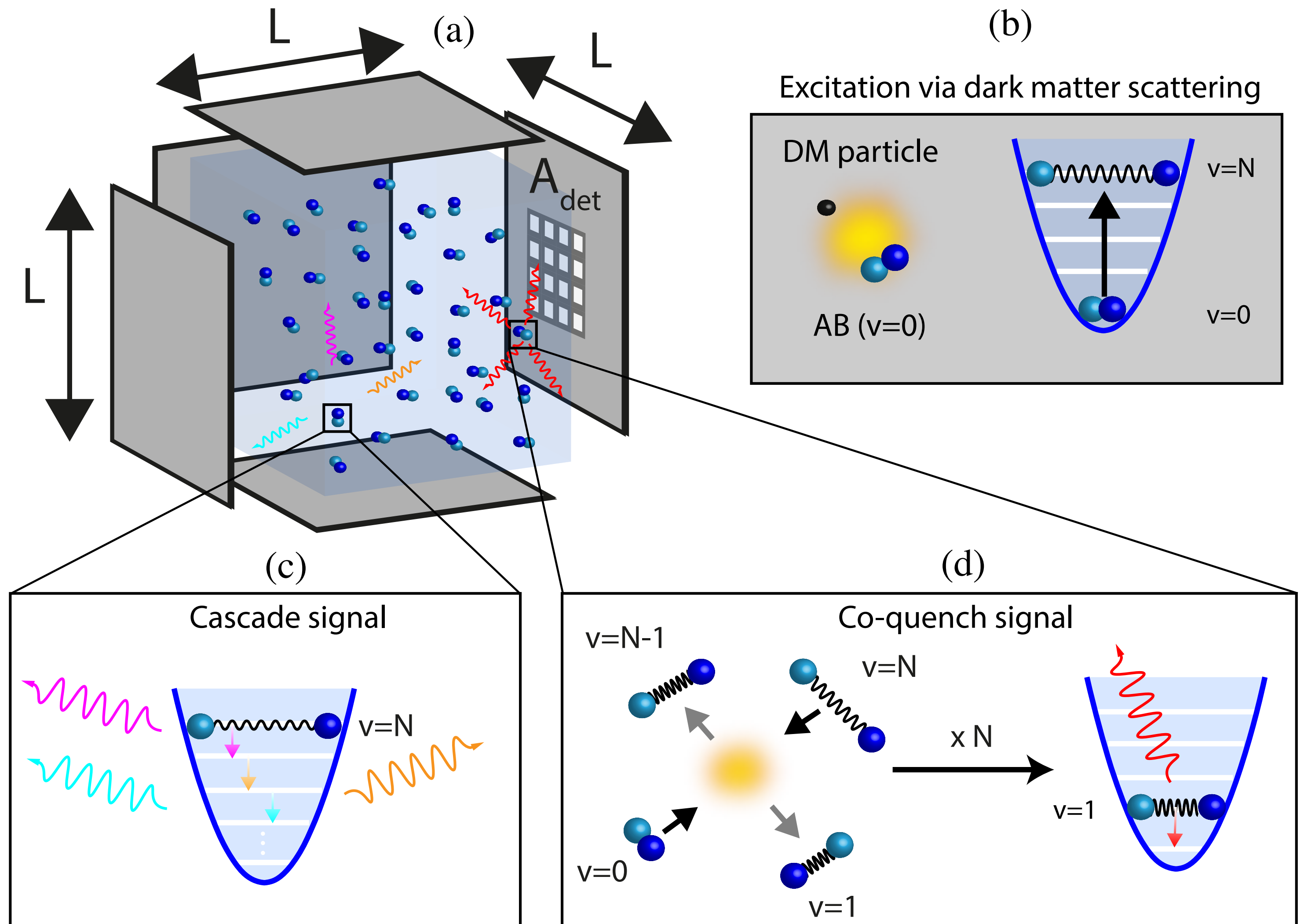
(d) Full detector





# Molecular Excitations and IR photon detection

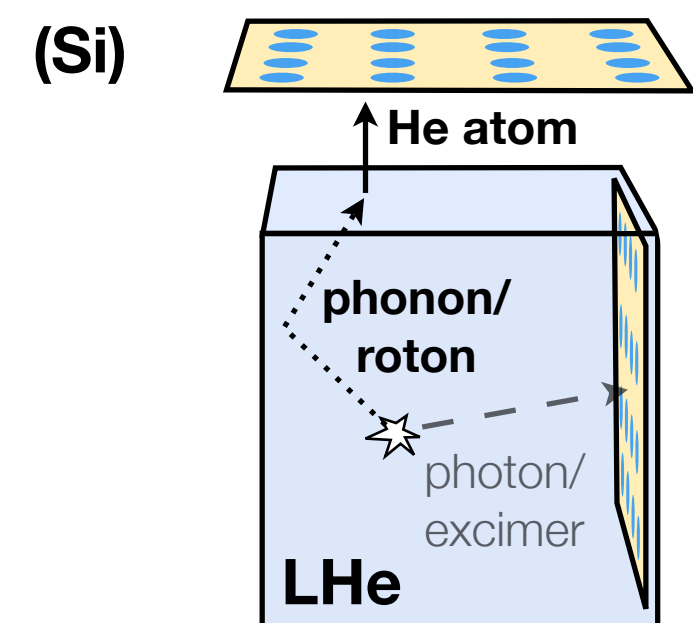
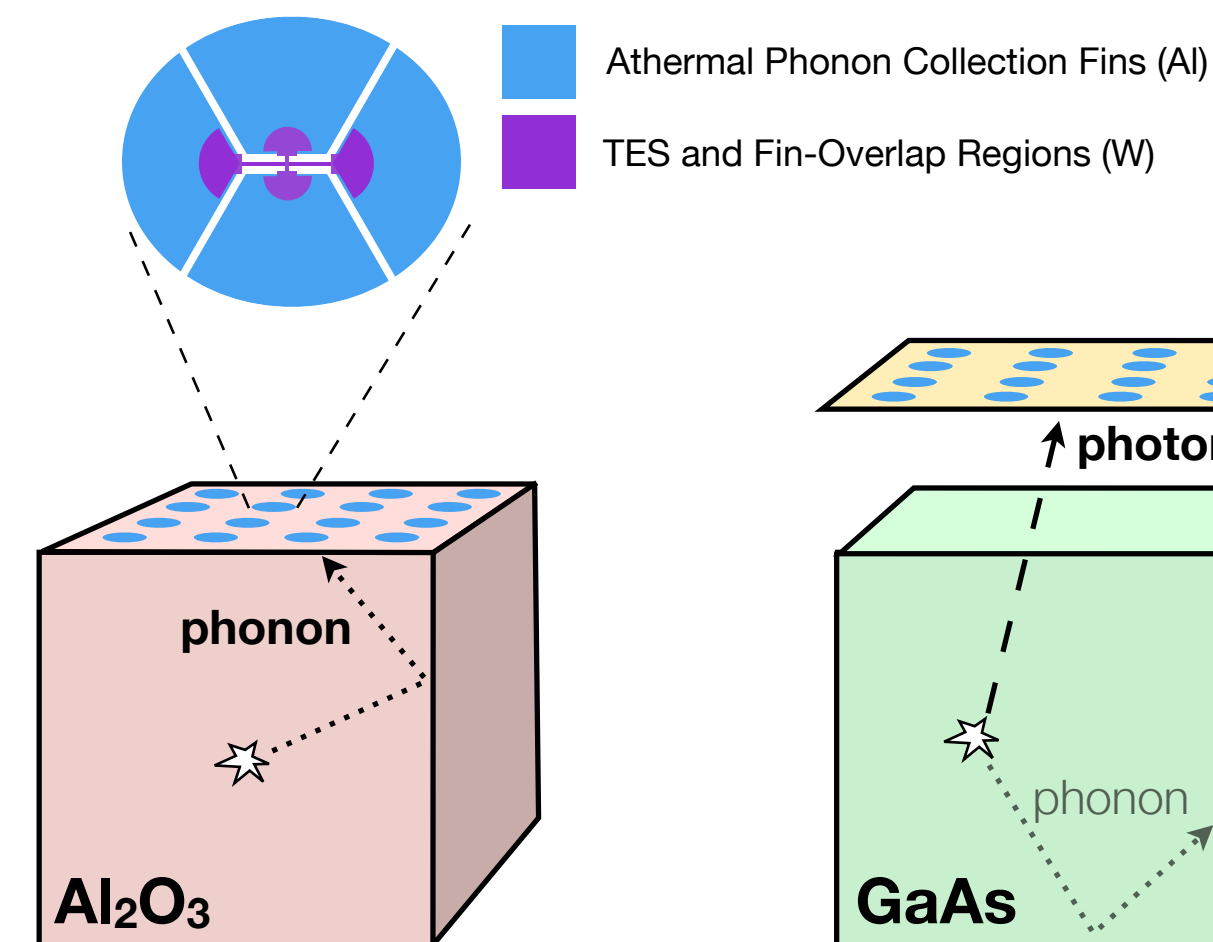
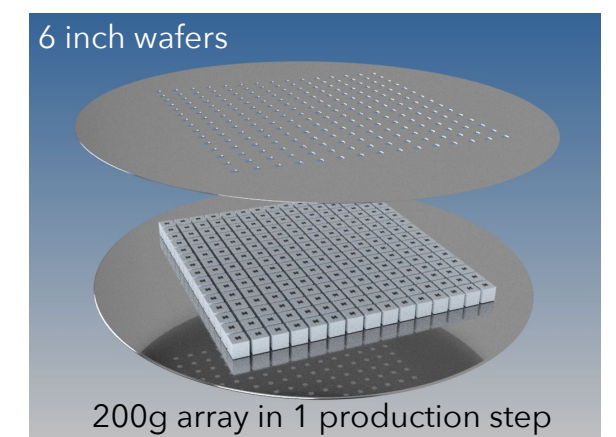
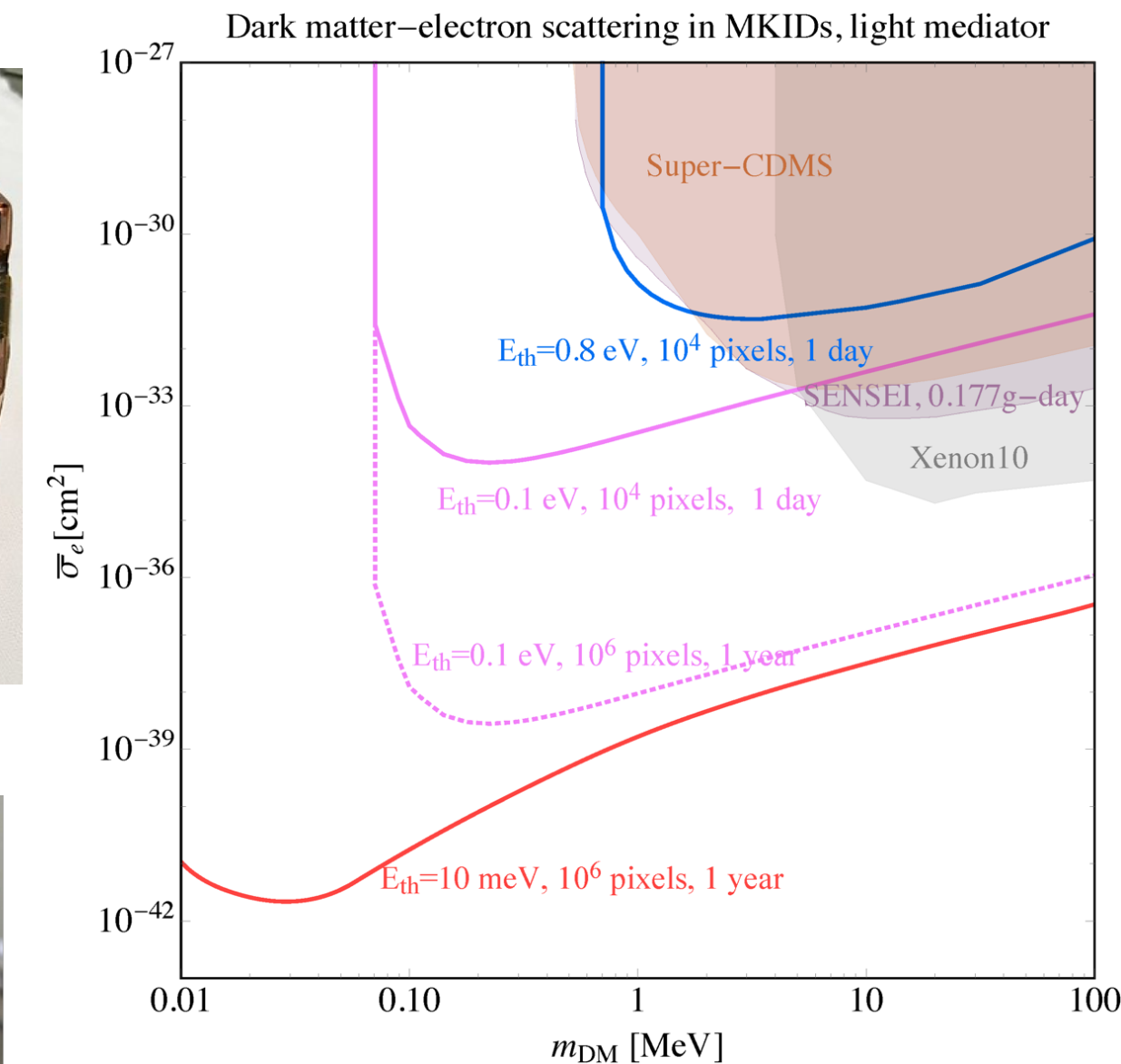
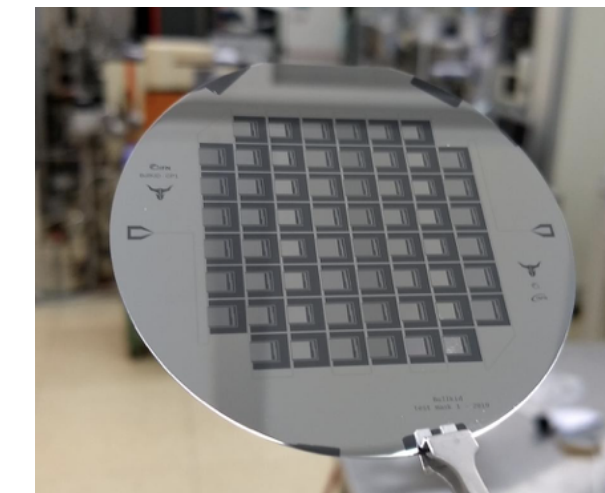
- Dark Matter interacts with low-pressure CO molecular target. Molecule goes to excited state, and relaxed by emitting an IR photon
- IR photon is sensed at surfaces with SNSPDs or other photon sensor.





# TES / KID / STJ / NTD / MagCals...

- Low-temperature phonon/quasiparticle detectors are being proposed by a large number of experiments, for dark matter, CMB, 0vBB, CEvNS, Cosmological Surveys, and other applications.
- R&D is needed in the following areas:
  - Lower thresholds
  - Lower noise readout
  - Multiplexing large number of detectors
  - Vibration, EMI, and other environmental noise reduction
  - Radioactive Background Mitigation
  - Other...





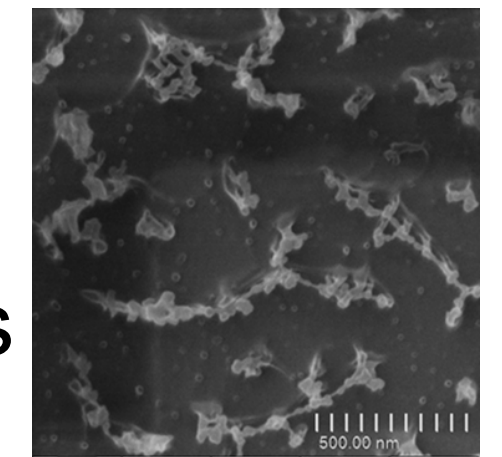
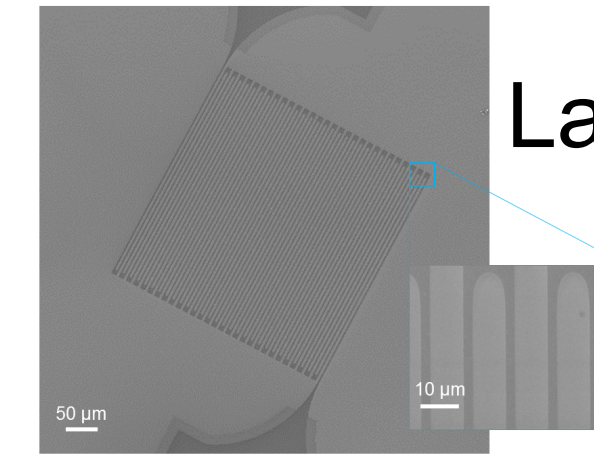
# SNSPD

## Superconducting Nanowire Single-Photon Detectors (SNSPDs)

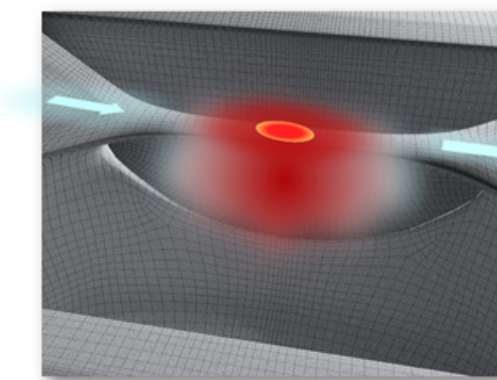
Overview: *Supercond. Sci. Technol.* 25 (2012) 063001

### Future prospects

Large-area, cm per side scale detectors

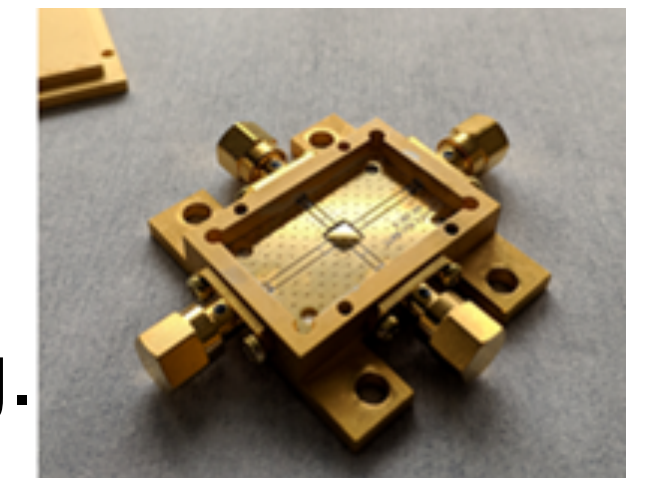


Detection in high- $T_c$  nanowires

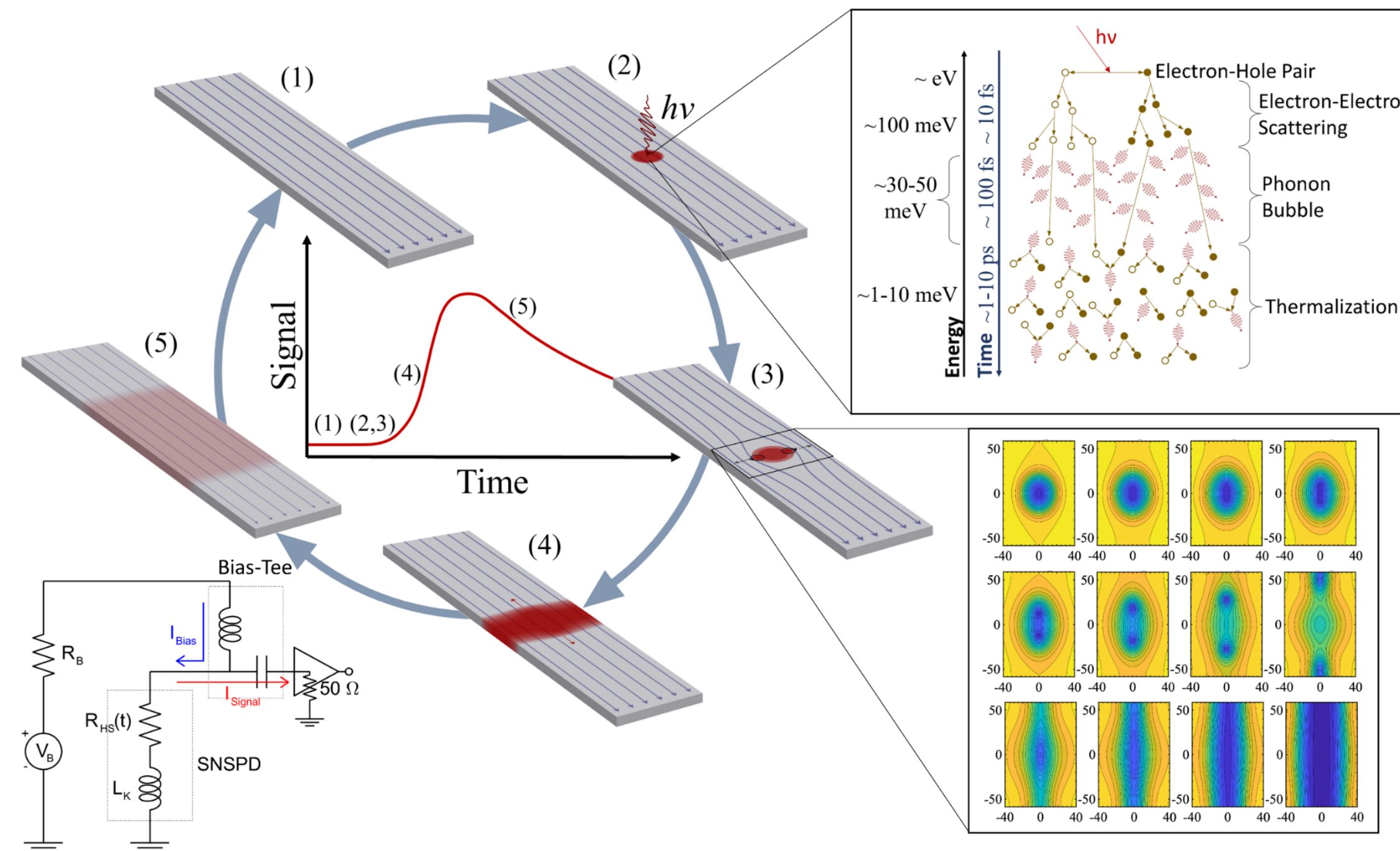


Mid-infrared (20  $\mu\text{m}$  and beyond) single-photon detection

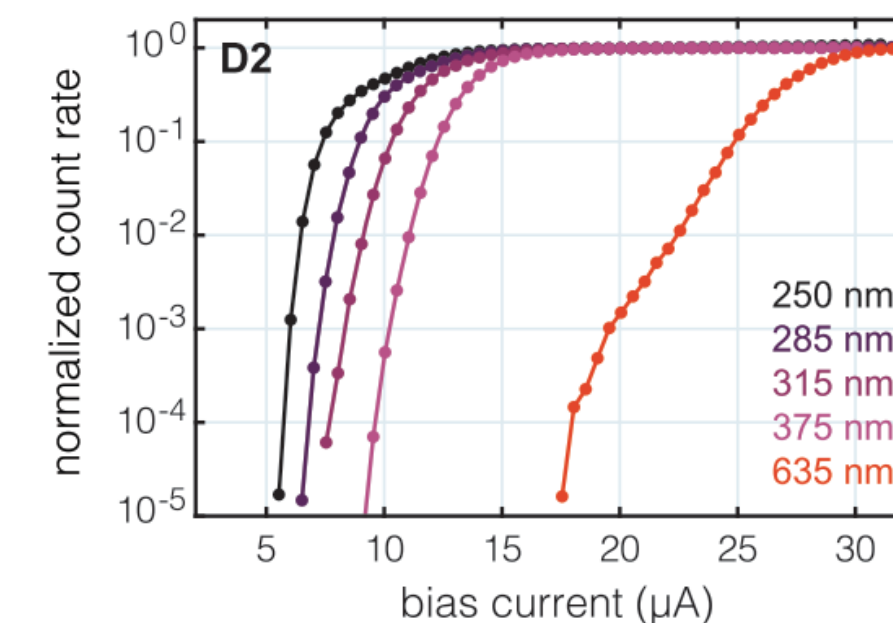
Multiplexed readout 100 Megapixel cameras and beyond. Integrated readout to preserve picosecond timing.



Extend high efficiency operation into Extreme-UV range.



- 98% efficiency demonstrated
- Good efficiency from UV to infrared range
- Count rate > 200 MHz
- ~ 1 dark count per day
- < 10 ps timing jitter

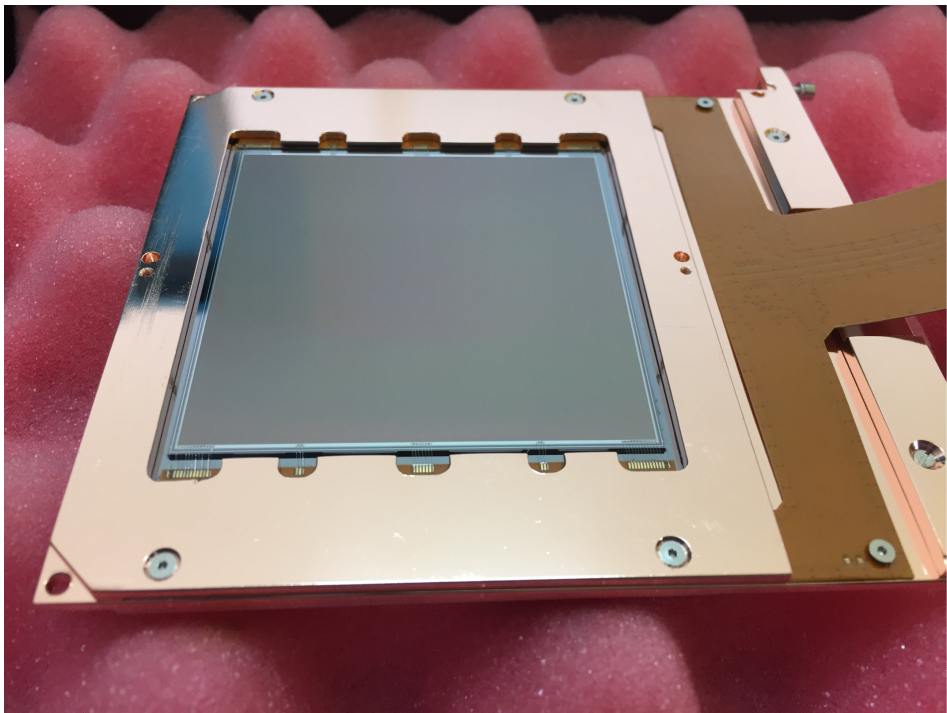
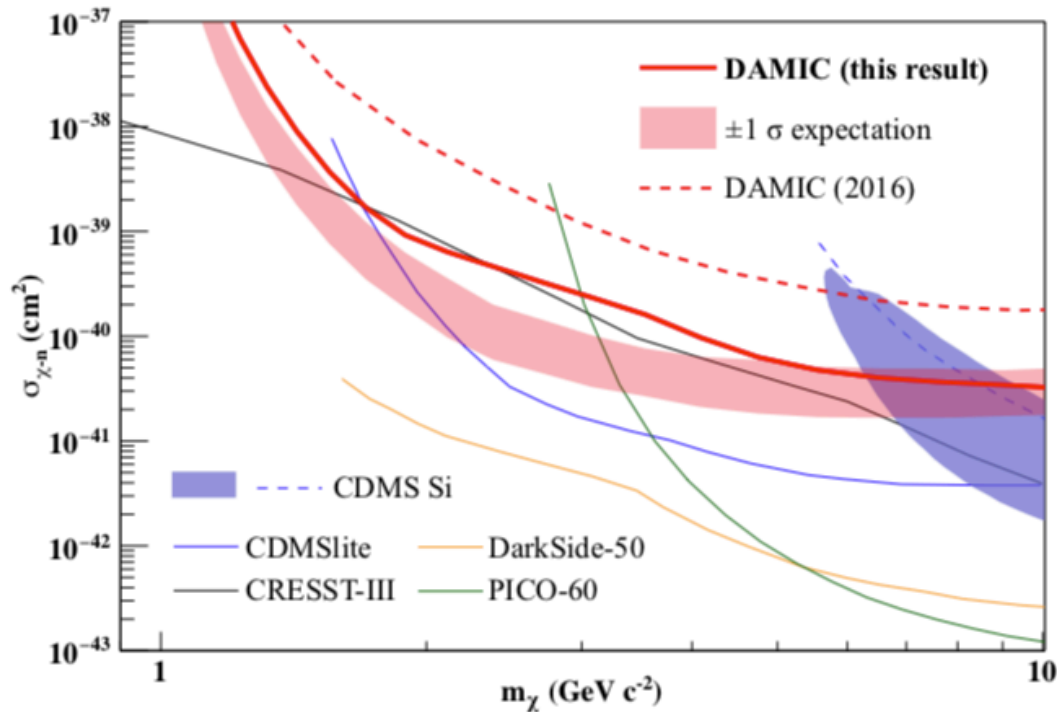




# CCDs for DM and Neutrino Physics

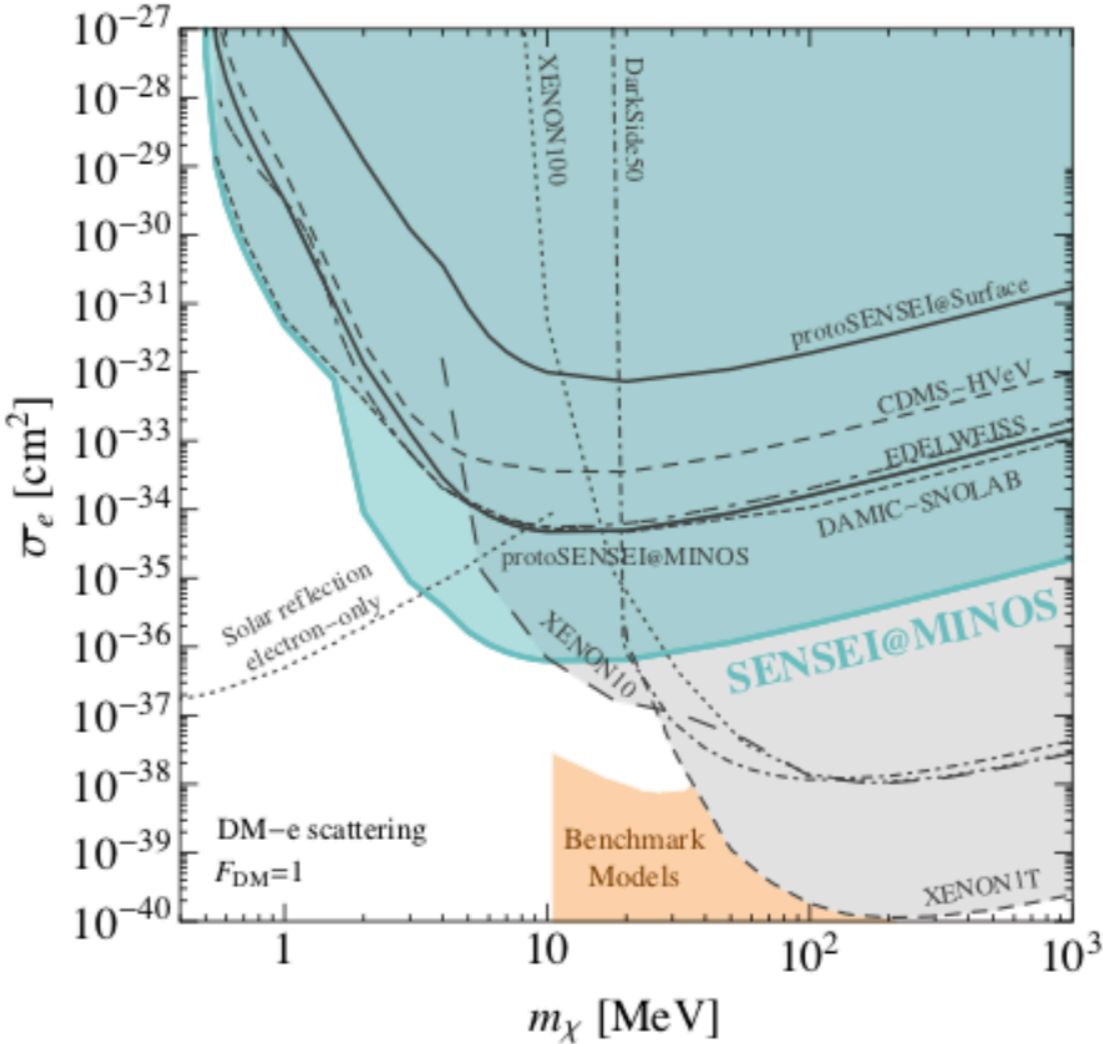
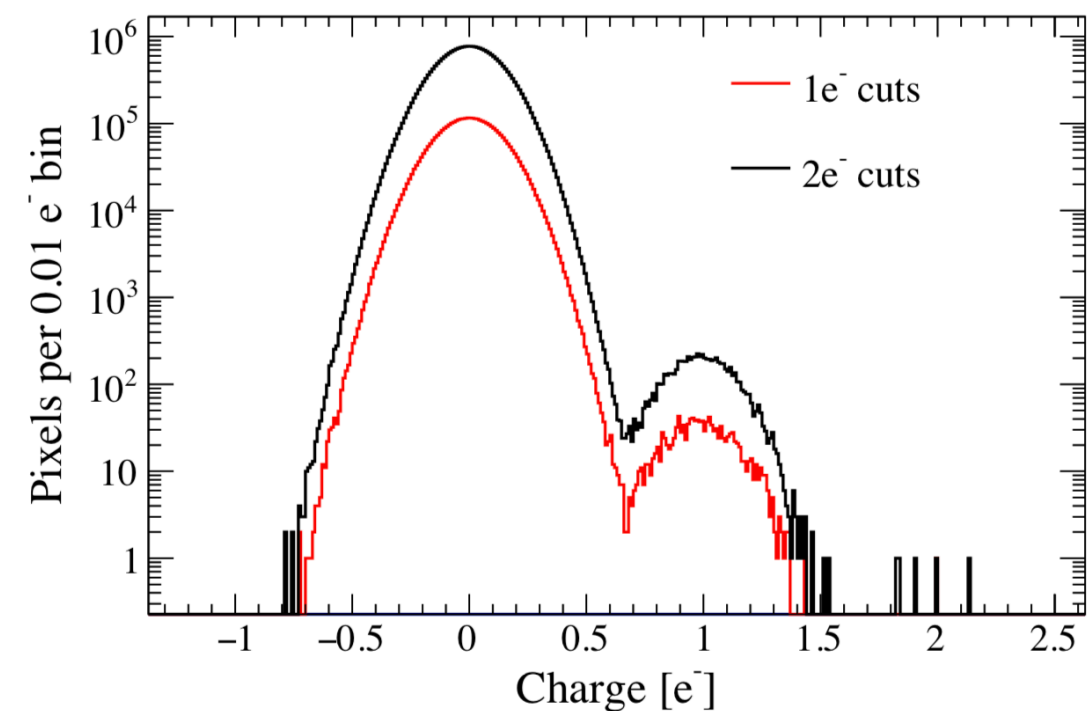
**DAMIC: ~50g at SNOLAB  
completed search for  
WIMPs with standard  
CCDs (~70eVee  
threshold).**

[arXiv:2007.15622](#)



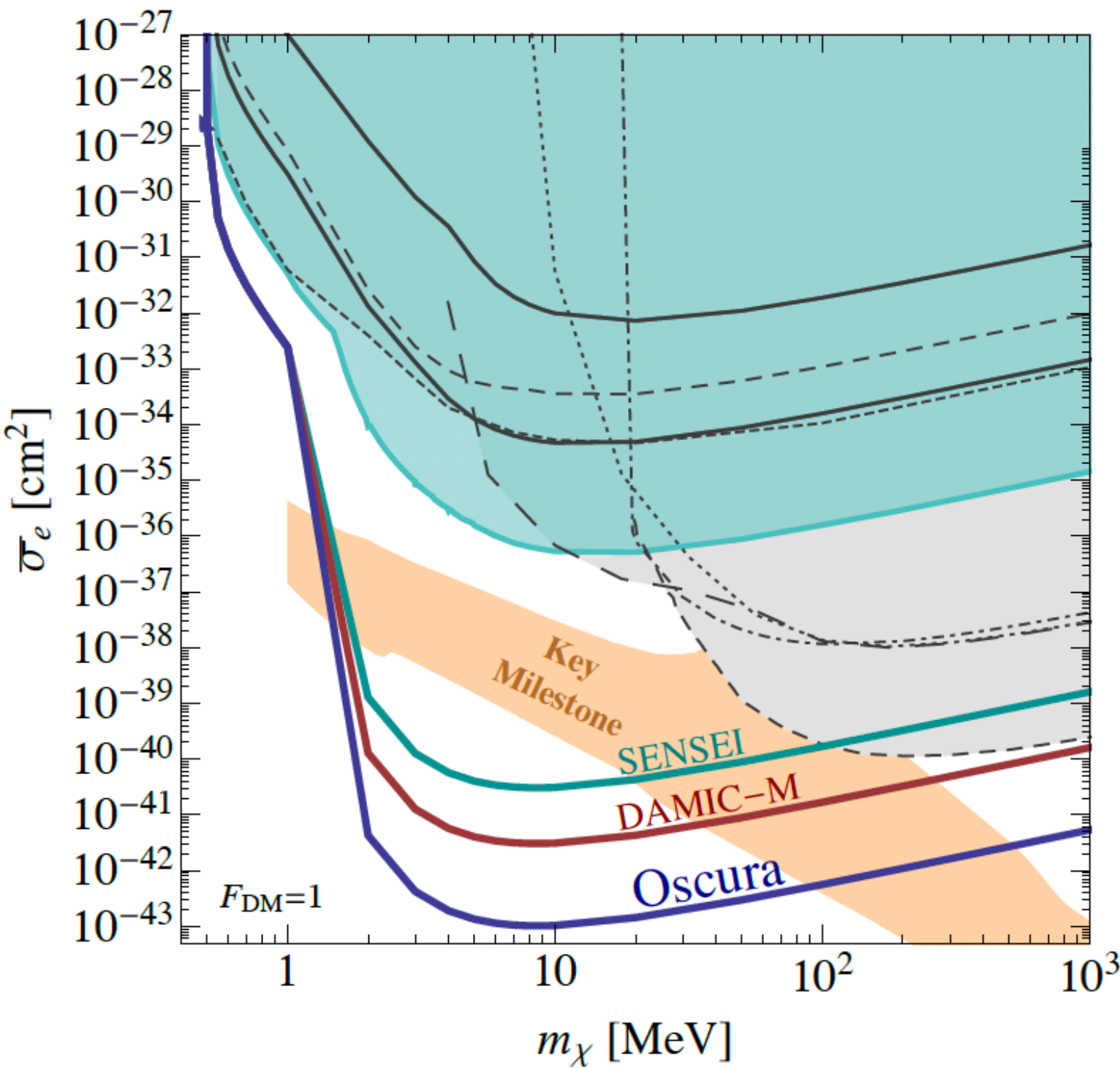
**Now: SENSEI ~2 gram  
detector producing  
world leading results at  
shallow underground  
site (MINOS).**

[arXiv:2004.11378](#)



**Future skipper-CCD projects:**  
2021(\*) : start SENSEI 100g @ SNOLAB  
2023(\*) : start DAMIC-M 1kg @ Modane  
2024(\*\*) : start building Oscura 10kg

(\*) fully funded  
(\*\*) R&D funded “New Ideas in Dark Matter”



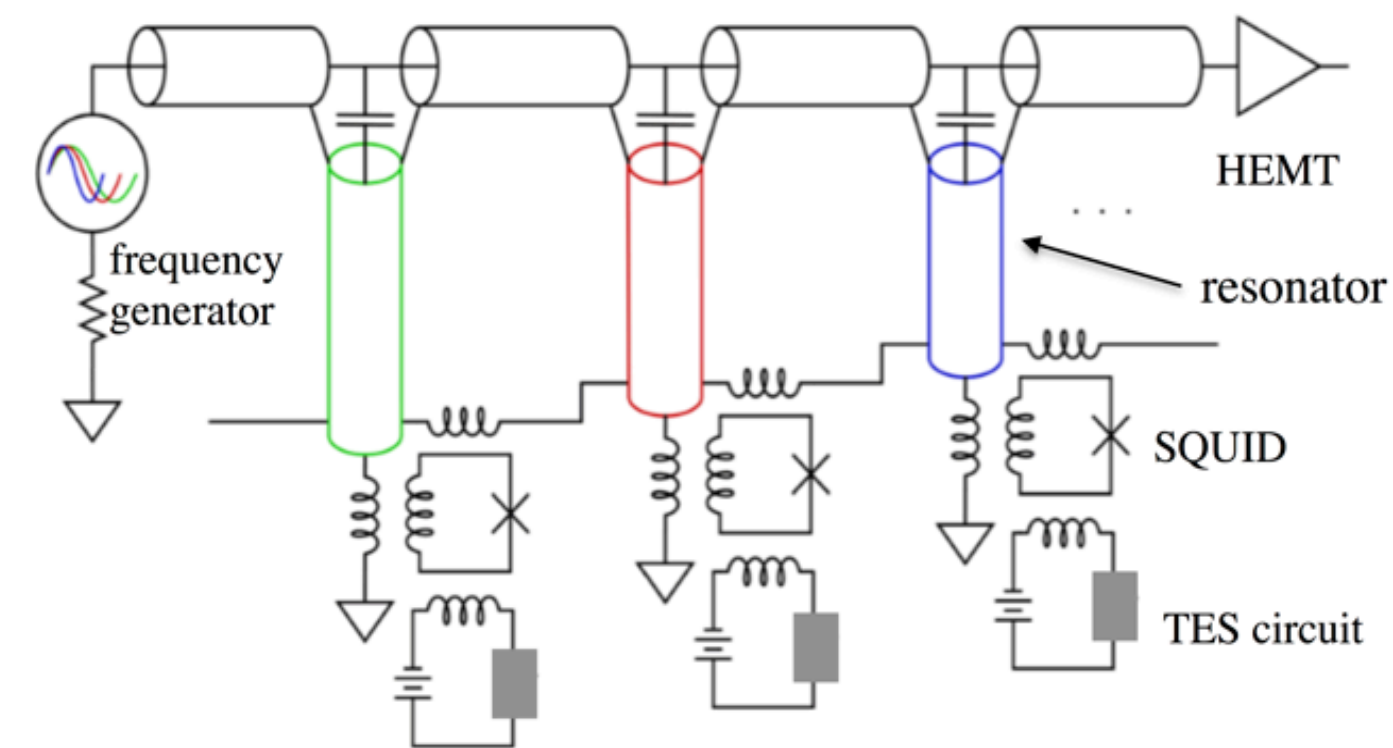
**a multi year plan, with a strong  
community support to take this  
technology to its full potential.**



# Readouts

- SQUID MUX: enable thousands of TES/MagCal readout channels in HEP experiments
- MKIDs/uMUX readouts: Lots of overlap between MKIDs and TES microwave MUX readouts in room temp components and DAC. Huge importance for future experiments
- Skipper CCD readout (for example using CMOS) for large mass next-gen experiments.

- Multiplexed readout for SNSPDs will allow large area detectors with 100 Megapixels and beyond.

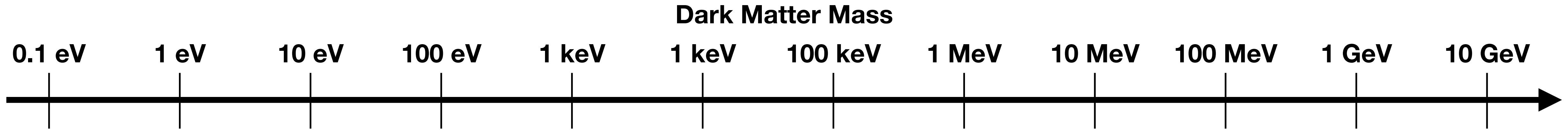


Microwave SQUID multiplexing

Mates et al. , Appl. Phys. Lett. **111**, 062601 (2017)

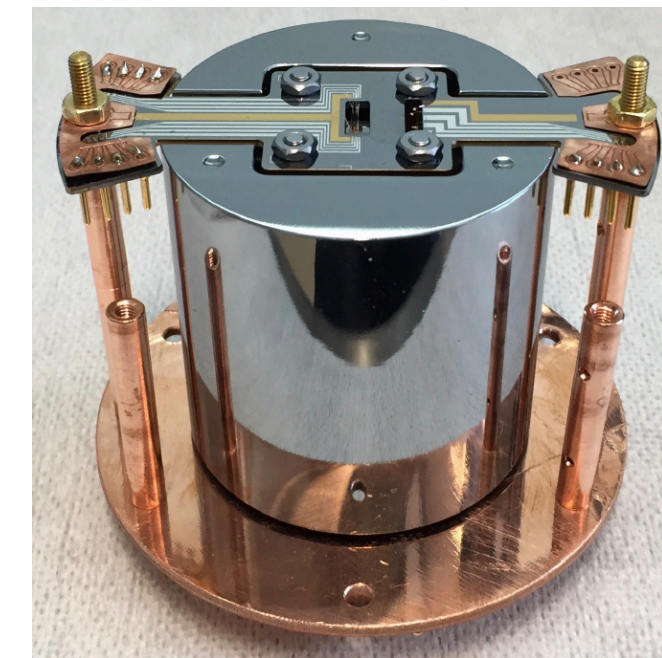


# Overlap with Neutrino Physics

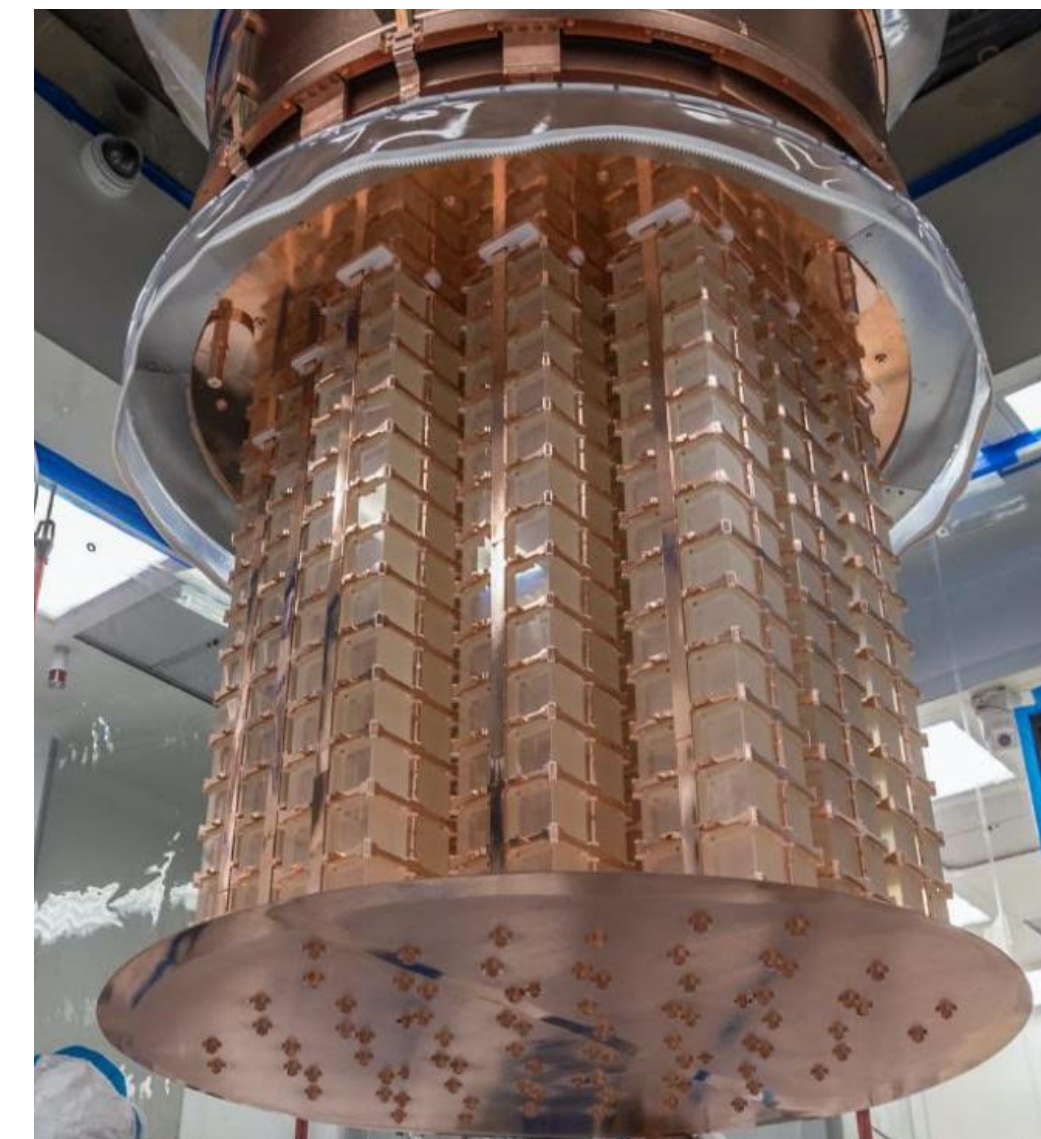
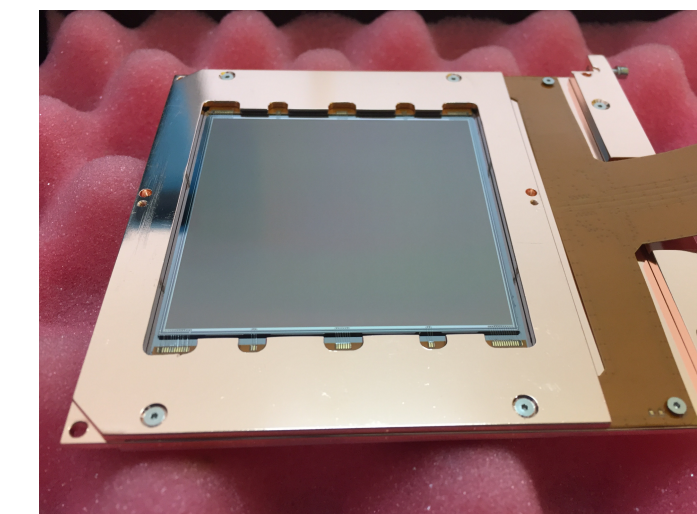


- Low-threshold technologies for light DM searches make excellent CEvNS detectors!
- Low-background large-mass arrays of detectors are needed both for future 0vBB and reaching the DM neutrino floor.
- Low-temperature techniques are shared by many of these experiments.
- Background discrimination and detector calibration are also shared amongst all three physics thrusts.

CEvNS Detectors



0vBB Decay Detectors





# Other Ideas

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- Theoretical work:
  - Modeling DM/EFT interactions, Quantum Computing algorithms
  - Modeling Quasiparticle transduction and lifetime; non-equilibrium dynamics
- Gravitational Wave detectors and Atomic clocks as DM detectors
- Materials Design for specific sensitivity to correlated quantum phenomena
- Mechanical quantum sensors with levitated mass arrays
- DM search with entangled electron spin qubits in  $^3\text{He}$
- Directional detection with solid state using quantum defects in diamond or carbon nanotubes



# Conclusions / Questions

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- There is a vibrant program of experiments covering the full mass range of particle dark matter.
- Many overlaps with QIS technologies, quantum sensing, axion-like searches.
- Neutrino detectors for both 0vBB and CEvNS have lots of technical overlap with many of these efforts.
- How do we balance technology R&D vs funding specific experiments?
- How can we best package the plethora of ideas into several strongly-motivated recommendations / white papers for our SNOWMASS report?